# Refinement of the Santamarina-Hfaiedh energy mesh between 22.5 eV and 11.4 keV

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# Abstract

We are investigating the possibility to improve the accuracy of lattice calculations at neutron energies where resonance self-shielding effects are important. The proposed improvement are related to a refinement of the Santamarina-Hfaiedh energy mesh (SHEM) between 22.5 eV and 11.14 keV, increasing the total number of energy groups from 281 to 361 groups. The self-shielded multigroup cross sections are obtained using a simplified and straightforward subgroup model, used in association with the 361–group SHEM. Using as many energy groups permits to avoid the explicit representation of correlated slowing-down effects. The resulting equations are becoming sufficiently simple to reintroduce an accurate representation of other physical effects that are generally neglected, namely the mutual shielding effect between different isotopes and the temperature correlation effect caused by an explicit temperature gradient in a resonant isotope. Therefore, the longstanding problems of resonance escape factor underestimation and Doppler coefficient are solved using the SPM subgroup method on the SHEM– 361 mesh. The resulting self-shielding model is shown to reach level of accuracies that are similar to those of a Monte-Carlo method.

#### 1. Introduction

The advent of advanced reactor designs and the availability of better computing resources are the main motivations for the development of more accurate lattice code algorithms. The selfshielding model is a good candidate for accuracy improvements. We are proposing the development of a new self-shielding method consistent with the introduction of a finer energy mesh, using as many as 361 energy groups. The proposed self-shielding method is based on a simplification of the Ribon extended model<sup>1</sup>,with the introduction of the capability to represent a temperature gradient in fuel. All the probability tables and weight correlated matrices are obtained with the CALENDF formalism.<sup>4</sup> The resulting formalism is referred as the subgroup projection method (SPM).<sup>6</sup>

The selected energy mesh, identified as

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SHEM-361, is an adaptation of the 281-group Santamarina-Hfaiedh energy mesh (SHEM) introduced in Ref. 2. We have reworked the energy domain between 22.5 eV and 11.14 keV, increasing the number of groups from 38 to 118. Using more energy groups is required in order to avoid the use of the slowing-down correlation model, as introduced in the Ribon extended method. The remaining SHEM limits above 11.14 keV are left unchanged. A try-and-error approach was used to optimize the choice of the energy limits. The application of the SPM is limited to groups 56 to 173, as most correlation effects vanishes above 11.14 keV. Below 22.5 eV, the correlation effects are important, but are taken into account directly by the optimized fine mesh of the original SHEM structure.

Modern lattice designs involve the introduction of different fuel pins with various isotopic contents, with or without burnable poisons. In some cases, the burnable poison is mixed with depleted fuel, leading to lower temperature pins. It is not always possible to find a unique effective temperature for all occurrences of a given isotope in the lattice, so that the accurate representation of temperature gradient effects in fuel cannot be avoided.

## 2. The subgroup projection method

The basic idea behind the SPM approach is to simplify the Ribon extended approach by removing the slowing-down correlation model and by introducing a new cross section correlation model compatible with any solution approach of the neutron transport equation. The removing of the slowing-down correlation model is possible by using as many as 118 energy groups in the domain between 22.5 eV and 11.14 keV. The SPM, as described in Ref. 6, is available in release 4.0.2 of DRAGON.<sup>9</sup> Open-source cross section libraries in SHEM–361 Draglib format are also available.

The computing properties of the proposed SPM can be summarized as follows:

- The SPM is a subgroup approach based on CALENDF-type probability tables.<sup>4</sup>
- The self-shielding calculations are limited to energies above 22.5 eV. The SPM is used below 11.14 keV and the ST approach with physical probability tables is used above.

- The self-shielding calculations are kept apart from the main flux calculation, thank to the Livolant-Jeanpierre factorization presented in Ref. 1. We observed that finemesh moderator and coolant spatial discretization can be avoided, leading to important CPU cost savings for the selfshielding calculations. We currently recommend to define a simplified geometry for self-shielding, different from the more complex geometry used for the main flux calculation.
- The subgroup equations are leading to independent solutions of the neutron transport equations that can be solved in parallel on multiple CPUs.
- The new cross section correlation model can effectively represent both *mutual shielding effects* and *temperature gradient effects* in fuel, *without* introducing additional CPU costs other than those associated with the use of 118 groups.
- Using as many as 118 energy groups between 22.5 eV and 11.14 keV, the SPH treatment of the self-shielded cross sections is not required. However, the SPH treatment is kept in remaining groups with a lethargy width greater than 0.1.
- The SPM can represent isotopic correlation effects in three different ways. A noncorrelation approximation is first available in case where resonances belonging to different isotopes are overlapping in a statistical way. A full-correlation model, similar to the model used in the ECCO lattice code, is available if different cross section sets are corresponding to the same isotope at different temperatures.<sup>5</sup> Finally, a general correlation model is available with the capability to represent any level of correlation, from no correlation to full correlation.

#### 3. Numerical results

We have based our validation study on a subset made of eight Rowlands pin-cell benchmark cases<sup>7</sup> and two original test-cases featuring a temperature gradient in fuel. The comparisons were made for light-water reactor pin-cells without leakage. Two types of pin-cell were studied, one UO2 fuelled (UOX), and the other UPuO2 fuelled, the latter in two versions with different isotopic compositions (MOX-1 and MOX-2). The effects of changes in temperature and water density were also calculated in order to examine the consistency of temperature calculation methods.

Five UOX cell cases were investigated. The first four are identical to those presented in Ref. 7. The fifth cell is an original PWR-HZP case, similar to case 3, with the arbitrary temperature profile of Table 1 in fuel. The goal if this test-case is to verify the numerical capabilities of the SPM to represent large temperature variations in a resonant isotope.

Table 1 Temperature gradient UOX rod.

| Outer radius | Mixture | Temperature |
|--------------|---------|-------------|
| (cm)         |         | (K)         |
| 0.2529822    | 1       | 1200.       |
| 0.3346640    | 1       | 1000.       |
| 0.3577709    | 1       | 800.        |
| 0.3794733    | 1       | 700.        |
| 0.3898718    | 1       | 650.        |
| 0.4          | 1       | 600.        |
| 0.45         | 2       | 600.        |
| 0.6770275    | 3       | 573.6       |

The five MOX cell Benchmarks are also based on simple Wigner-Seitz unit cells. We investigated the effects of an increase in fuel temperature with two different isotopic vectors. Again, a fifth MOX cell was defined with the arbitrary temperature profile of Table 2 in fuel.

Table 2 Temperature gradient in MOX rod.

| Outer radius | Mixture | Temperature |
|--------------|---------|-------------|
| (cm)         |         | (K)         |
| 0.259307     | 1       | 900.        |
| 0.343031     | 1       | 700.        |
| 0.366715     | 1       | 500.        |
| 0.388960     | 1       | 400.        |
| 0.3996186    | 1       | 350.        |
| 0.41         | 1       | 300.        |
| 0.475        | 2       | 300.        |
| 0.710879     | 3       | 293.6       |

The one-neutron source validation tests are limited to the resolved energy domain where it is possible to precisely define the resonant cross sections. The scattering kernel is assumed to be purely elastic. Cross sections were defined in the resolved energy domain and distributed over SHEM–361 energy groups 56 to 173, located between 22.5 eV and 11.14 keV. A 1.0 n/cm3/s source was placed in group number 56, located between 9.1188 keV and 11.138 keV and the absorption rates are computed in the remaining energy groups.

The UOX and MOX calculations use the same multigroup cross-section library based on the JEF 2.2 evaluation. Cross section libraries in PENDF and Draglib formats were build from scratch with NJOY release 99.259+upnea027. We selected the modified 361–group SHEM mesh presented in this study, with the correlation model set for all resonant isotopes. Draglib–formatted data is including temperature–dependent Autolib data for all resonant isotopes between 22.5 eV and 11.14 keV. The elementary lethargy width of the Autolib data is  $5 \times 10^{-4}$ .

Two computer codes have been used to perform these tests:

- 1. A computer code, named CESCOL, makes it possible to solve a fixed-source slowingdown equation using an elastic slowingdown operator for a mixture of heavy (resonant) isotopes in the resolved energy domain.<sup>8</sup> Heterogeneous cases can also be treated using collision probability (CP) techniques and used to generate reference solutions.
- 2. A self-shielding operator was written in the DRAGON Version4 lattice code<sup>9</sup> based on the SPM subgroup model. Self-shielded cross sections are obtained for a coarse energy grid and used in the existing CP flux solution operators. Consistency is emphasized by using the same CP calculation operator in both heterogeneous CESCOL and lattice code calculations.

We studied the absorption rates for the resonant isotopes in energy groups 56 to 173 and reported the discrepancies between CESCOL and lattice code calculations for every UOX benchmark.

The main purpose of the numerical tests was to compare the proposed self-shielding methodology with reference CESCOL calculations. The corresponding numerical results are presented in Tables 3 and 4 for various benchmark conditions. We are reporting global error values for maximum  $\epsilon^{\max}$ , averaged  $\bar{\epsilon}$  and integrated error  $\epsilon^{int}$  isotopic and spatially-dependent  $\epsilon^{int}$  values. The percent errors on absorption rates are plotted in Figs. 1 and 2. Note that the presence of the 1.0 n/cm<sup>3</sup>/s source produces unphysical transients on absorption rates in groups 56 and 57.

An important observation is related to the use of the SPM-type general correlation model. In UOX cases, this model can be avoided in all cases, provided that the temperature gradients are treated with the ECCO-type full-correlation model. Neglecting the temperature correlation effects in the fifth test-case leads to a maximum error  $\epsilon^{\max}$  reaching 84% in this case. In MOX cases, the correlation effects are more important. The application of the SPM–type general correlation model permits to reduce the maximum error below 4% and to divide by a factor of two the maximum error in each of the five MOX testcases. It is important to note that the correlation model introduces no additional CPU costs, so that it can be left active in all situations. Both ECCO– and SPM–type correlation models permit to treat correctly the fuel temperature coefficient in presence of temperature gradient.

Table 3

Summary of UOX one-neutron source benchmarks. General correlation model is active.

|   | Case 1          | Case 2         | Case 3       | Case 4            | Case 5           |
|---|-----------------|----------------|--------------|-------------------|------------------|
|   | Isothermal 293K | Reduced $H_2O$ | Fuel at 900K | Isothermal $574K$ | Temperature      |
|   |                 | density        |              |                   | gradient in fuel |
| $\epsilon^{\text{int}}$ (%)                     | -0.059          | -0.145         | 0.057        | -0.014            | 0.070            |
| $\bar{\epsilon}$ (%)                            | 0.546           | 0.555          | 0.519        | 0.542             | 0.529            |
| $\epsilon^{\max}$ (%)                           | 3.337           | 2.781          | 2.513        | 2.602             | 2.419            |
| in group  | 142             | 142            | 106          | 106               | 79               |
| $^{235}$ U $\epsilon^{int}$ (%)                 | 0.353           | 0.361          | 0.360        | 0.351             | 0.362            |
| $^{238}\mathrm{U}~\epsilon^{\mathrm{int}}~(\%)$ | -0.242          | -0.377         | -0.066       | -0.170            | -0.048           |
| $^{238}$ U $\epsilon^{int}$ (%)                 |                 |                |              |                   |                  |
| shell 1   | 0.051           | -0.117         | 0.357        | 0.201             | 0.520            |
| shell 2<br>shell 3<br>shell 4                   | 0.023           | -0.144         | 0.353        | 0.218             | 0.419            |
|   | -0.299          | -0.425         | 0.090        | -0.085            | -0.090           |
|   | -0.508          | -0.611         | -0.378       | -0.515            | -0.645           |
| shell 5   | -0.806          | -0.884         | -1.129       | -1.095            | -1.347           |
| shell 6   | -1.158          | -1.262         | -1.669       | -1.540            | -1.919           |

|  | MOX fuel 1      | MOX fuel 1   | MOX fuel 2      | MOX fuel 2   | MOX fuel 1       |
|--|-----------------|--------------|-----------------|--------------|------------------|
|  | Isothermal 300K | Fuel at 560K | Isothermal 300K | Fuel at 560K | Temperature      |
|  |                 |              |                 |              | gradient in fuel |
| $\epsilon^{\text{int}}$ (%)                      | 0.098           | 0.157        | 0.140           | 0.211        | 0.211            |
| $\bar{\epsilon}$ (%)                             | 0.751           | 0.777        | 0.816           | 0.841        | 0.770            |
| $\epsilon^{\max}$ (%)                            | 3.079           | 3.458        | 3.495           | 3.508        | 3.318            |
| in group   | 124             | 124          | 67              | 124          | 124              |
| $^{235}$ U $\epsilon^{int}$ (%)                  | 0.420           | 0.402        | 0.403           | 0.382        | 0.389            |
| $^{238}$ U $\epsilon^{int}$ (%)                  | -0.209          | -0.024       | -0.204          | -0.013       | 0.121            |
| $^{238}\mathrm{Pu}\;\epsilon^{\mathrm{int}}$ (%) | 0.237           | 0.303        | 0.355           | 0.421        | 0.343            |
| $^{239}$ Pu $\epsilon^{\text{int}}$ (%)          | 0.241           | 0.220        | 0.380           | 0.372        | 0.198            |
| $^{240}$ Pu $\epsilon^{\text{int}}$ (%)          | 0.597           | 0.536        | 0.504           | 0.465        | 0.569            |
| $^{241}$ Pu $\epsilon^{\text{int}}$ (%)          | 0.372           | 0.363        | 0.345           | 0.336        | 0.351            |
| $^{242}$ Pu $\epsilon^{int}$ (%)                 | 0.657           | 0.400        | 0.464           | 0.278        | 0.206            |
| <sup>241</sup> Am $\epsilon^{\text{int}}$ (%)    | 0.342           | 0.335        | 0.324           | 0.313        | 0.317            |
| $^{238}$ U $\epsilon^{int}$ (%)                  |                 |              |                 |              |                  |
| shell 1  | 0.242           | 0.557        | 0.248           | 0.564        | 0.860            |
| shell 2  | 0.079           | 0.401        | 0.089           | 0.410        | 0.530            |
| shell 3  | -0.309          | -0.051       | -0.301          | -0.036       | -0.350           |
| shell 4  | -0.570          | -0.548       | -0.571          | -0.531       | -0.743           |
| shell 5  | -0.946          | -1.160       | -0.946          | -1.147       | -0.770           |
| shell 6  | -1.513          | -1.696       | -1.513          | -1.685       | -1.853           |

Summary of MOX one-neutron source benchmarks. General correlation model is active.

## Conclusions

Table 4

The 361-group refined Santamarina-Hfaiedh energy mesh is permitting a better representation of self-shielding phenomena between 22.5 eV and 11.14 keV. The subgroup projection method is a good candidate for performing resonance selfshielding calculations in association with the refined energy mesh. It involves twice more energy groups than XMAS-172g mesh used in thermal reactors, but it allows high levels of accuracy. Furthermore, this optimized SHEM-361g could be used in FBR calculations in order to reduce drastically the computing time linked to the current 1968 group structure. The subgroup projection method is compatible with any type of solution of the transport equation. It permits the representation of distributed self-shielding effects, mutual shielding effects and temperature gradient effects. Particularly, it solves the longstanding problems of resonance escape factor and Doppler coefficient calculations in MOX assemblies and HCLWR lattices.

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Fig. 1 Percent error on absorption rates. UOX cases. General correlation model is active.

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Fig. 2 Percent error on absorption rates. MOX cases. General correlation model is active.

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# A The 361–group Santamarina-Hfaiedh mesh

The 361–group Santamarina-Hfaiedh mesh (SHEM–361) is based on the original SHEM–281 structure presented in Refs. 2 and 3. We have reworked the energy limits between 22.5 eV and

11.14 keV, increasing the number of groups in this domain from 38 to 118. The SPM is applied in groups 56 to 173 (included), corresponding to the resolved resonant energy groups. The improved structure of the SHEM–361 mesh is given in the following tables.



| 93  | 10.466 | 2.8489E + 02 | 0.030 |   |
|-----|--------|--------------|-------|---|
| 94  | 10.496 | 2.7647E + 02 | 0.030 |   |
| 95  | 10.526 | 2.6830E + 02 | 0.044 |   |
| 96  | 10.570 | 2.5675E + 02 | 0.060 |   |
| 97  | 10.630 | 2.4180E + 02 | 0.026 | ← Pu240 (239.3 eV), U238 (237.4 eV)           |
| 98  | 10.656 | 2.3559E + 02 | 0.049 |   |
| 99  | 10.705 | 2.2432E + 02 | 0.056 |   |
| 100 | 10.761 | 2.1211E + 02 | 0.054 |   |
| 101 | 10.815 | 2.0096E + 02 | 0.025 |   |
| 102 | 10.840 | 1.9600E + 02 | 0.015 |   |
| 103 | 10.855 | 1.9308E + 02 | 0.015 |   |
| 104 | 10.870 | 1.9020E + 02 | 0.007 | ← U238 (189.7 eV)                             |
| 105 | 10.877 | 1.8888E + 02 | 0.007 |   |
| 106 | 10.884 | 1.8756E + 02 | 0.007 |   |
| 107 | 10.891 | 1.8625E + 02 | 0.007 | ← Pu240 (185.8 eV)                            |
| 108 | 10.898 | 1.8495E + 02 | 0.009 |   |
| 109 | 10.907 | 1.8329E + 02 | 0.045 |   |
| 110 | 10.952 | 1.7523E + 02 | 0.045 | ← Pu240 (170.1 eV)                            |
| 111 | 10.997 | 1.6752E + 02 | 0.027 | ← U238 (165.3 eV)                             |
| 112 | 11.024 | 1.6306E + 02 | 0.056 | ← Pu240 (162.7 eV)                            |
| 113 | 11.080 | 1.5418E + 02 | 0.050 | ← Pu240 (151.9 eV)                            |
| 114 | 11.130 | 1.4666E + 02 | 0.050 |   |
| 115 | 11.180 | 1.3950E + 02 | 0.050 | ← Pu240 (135.3 eV)                            |
| 116 | 11.230 | 1.3270E + 02 | 0.050 |   |
| 117 | 11.280 | 1.2623E + 02 | 0.046 | ← Pu240 (121.7 eV)                            |
| 118 | 11.326 | 1.2055E + 02 | 0.025 |   |
| 119 | 11.351 | 1.1758E + 02 | 0.009 | ← U238 (116.9 eV)                             |
| 120 | 11.360 | 1.1652E + 02 | 0.009 |   |
| 121 | 11.369 | 1.1548E + 02 | 0.023 |   |
| 122 | 11.392 | 1.1285E + 02 | 0.023 |   |
| 123 | 11.415 | 1.1029E + 02 | 0.043 |   |
| 124 | 11.458 | 1.0565E + 02 | 0.025 | - Pu240 (105.1 eV)                            |
| 125 | 11.483 | 1.0304E + 02 | 0.009 | ← U238 (102.6 eV)                             |
| 126 | 11.492 | 1.0211E + 02 | 0.005 |   |
| 127 | 11.497 | 1.0161E + 02 | 0.005 |   |
| 128 | 11.502 | 1.0110E + 02 | 0.005 |   |
| 129 | 11.507 | 1.0059E + 02 | 0.033 |   |
| 130 | 11.540 | 9.7329E + 01 | 0.042 |   |
| 131 | 11.582 | 9.3326E + 01 | 0.050 | $\leftarrow$ Pu240 (90.8 eV), Pu240 (92.5 eV) |
| 132 | 11.632 | 8.8774E + 01 | 0.056 |   |
| 133 | 11.688 | 8.3939E + 01 | 0.056 | $\leftarrow$ U238 (80.7 eV)                   |
| 134 | 11.744 | 7.9368E + 01 | 0.039 |   |
| 135 | 11.783 | 7.6332E + 01 | 0.037 |   |
| 136 | 11.820 | 7.3559E + 01 | 0.023 | ← Pu240 (72.8 eV)                             |
| 137 | 11.843 | 7.1887E + 01 | 0.040 |   |
|     |        |              |       |   |

| 138 | 11.883 | 6.9068E + 01   | 0.033 |                                    |
|-----|--------|----------------|-------|------------------------------------|
| 139 | 11.916 | 6.6826E + 01   | 0.005 | ← Pu240 (66.6 eV)                  |
| 140 | 11.921 | 6.6493E + 01   | 0.005 |                                    |
| 141 | 11.926 | 6.6161E + 01   | 0.005 | ← U238 (66.0 eV), Pu239 (65.9 eV)  |
| 142 | 11.931 | $6.5831E{+}01$ | 0.005 |                                    |
| 143 | 11.936 | $6.5503E{+}01$ | 0.007 |                                    |
| 144 | 11.943 | 6.5046E + 01   | 0.007 |                                    |
| 145 | 11.950 | $6.4592E{+}01$ | 0.015 |                                    |
| 146 | 11.965 | $6.3631E{+}01$ | 0.021 |                                    |
| 147 | 11.986 | 6.2308E + 01   | 0.039 |                                    |
| 148 | 12.025 | $5.9925E{+}01$ | 0.049 |                                    |
| 149 | 12.074 | $5.7059E{+}01$ | 0.054 |                                    |
| 150 | 12.128 | 5.4060E + 01   | 0.020 | ← Pu242 (53.46 eV)                 |
| 151 | 12.148 | 5.2990E + 01   | 0.023 |                                    |
| 152 | 12.171 | 5.1785E + 01   | 0.050 |                                    |
| 153 | 12.221 | $4.9259E{+}01$ | 0.036 | ← Hf176 (48.3 eV), Hf177 (48.9 eV) |
| 154 | 12.257 | 4.7517E + 01   | 0.028 |                                    |
| 155 | 12.285 | $4.6205E{+}01$ | 0.020 |                                    |
| 156 | 12.305 | $4.5290E{+}01$ | 0.025 | - Mo95 (44.9 eV)                   |
| 157 | 12.330 | 4.4172E + 01   | 0.024 |                                    |
| 158 | 12.354 | 4.3125E + 01   | 0.023 | $\leftarrow$ Nd145 (42.5 eV)       |
| 159 | 12.377 | 4.2144E + 01   | 0.022 | ← Pu240 (41.6 eV)                  |
| 160 | 12.399 | 4.1227E + 01   | 0.037 | ← Ag109 (40.1 eV)                  |
| 161 | 12.436 | 3.9730E + 01   | 0.024 | $\leftarrow$ In115 (39.6 eV)       |
| 162 | 12.460 | 3.8787E + 01   | 0.026 | ← Pu240 (38.3 eV)                  |
| 163 | 12.486 | 3.7792E + 01   | 0.013 | -                                  |
| 164 | 12.499 | 3.7304E + 01   | 0.012 |                                    |
| 165 | 12.511 | 3.6859E + 01   | 0.012 | ← U238 (36.7 eV), Hf179 (36.6 eV)  |
| 166 | 12.523 | 3.6419E + 01   | 0.010 | $\mapsto$ Hf177 (36.1 eV. 37.0 eV) |
| 167 | 12.533 | 3.6057E + 01   | 0.010 |                                    |
| 168 | 12.543 | 3.5698E + 01   | 0.033 |                                    |
| 169 | 12.576 | $3.4539E{+}01$ | 0.043 |                                    |
| 170 | 12.619 | 3.3085E + 01   | 0.043 |                                    |
| 171 | 12.662 | 3.1693E + 01   | 0.128 |                                    |
| 172 | 12.790 | 2.7885E + 01   | 0.123 |                                    |
| 173 | 12.913 | 2.4658E + 01   | 0.090 | $\leftarrow$ U235 (23.42 eV)       |